

Status of the MEG Experiment

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for the MEG collaboration

Outline

- Physics motivations for the MEG experiment
- MEG detector
- Status of the sub-detectors
 - ◆ Beam line
 - ◆ Photon detector
 - ◆ Positron spectrometer
 - Magnet
 - Drift chamber
 - Timing counter
 - ◆ Trigger, DAQ, and slow control
- Summary

MEG Collaboration

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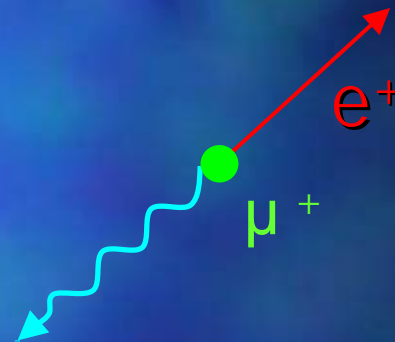
⁶ PSI Villigen, Switzerland

⁷ University of Tokyo, Tokyo, Japan

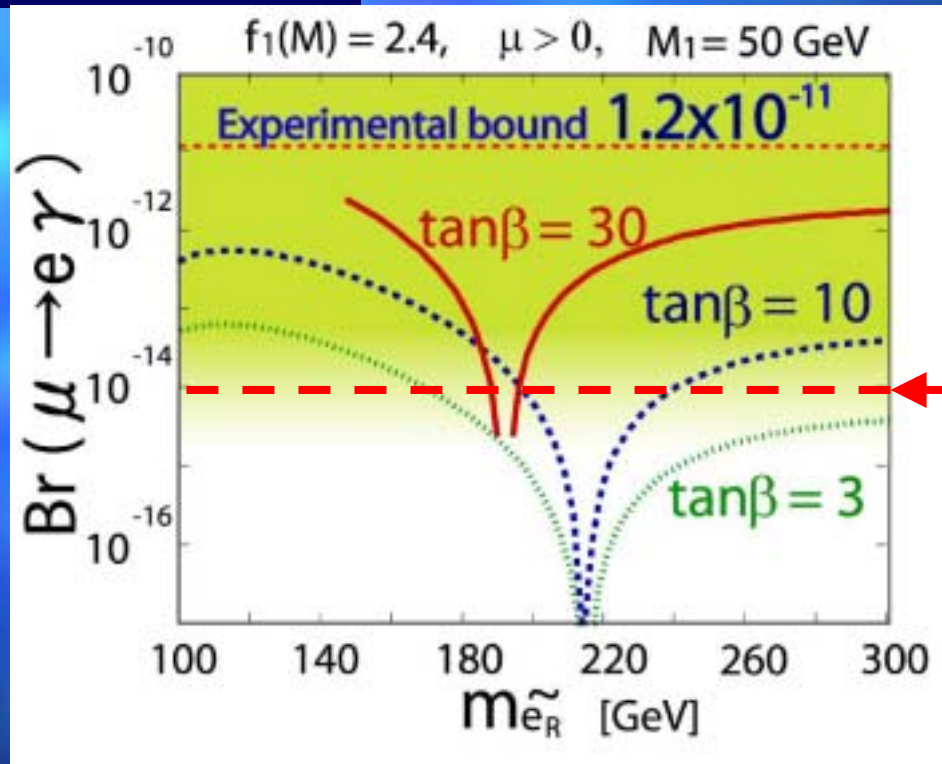
⁸ Waseda University, Tokyo, Japan



- Event signature
 - Back to back
 - Time coincident
 - $E_e = E_\gamma = 52.8\text{MeV}$
- Lepton-family-number nonconserving process
- Extremely small branching ratio in the standard model with finite neutrino mass
 - ex.) $\text{BR}(\mu \rightarrow e \gamma) \sim 10^{-52}$ for $m_\nu \sim 0.05\text{eV}$
- Sensitive to physics beyond the standard model
 - SUSY-GUT, SUSY+ R , ...
- Present experimental bound
 - $\text{BR}(\mu^+ \rightarrow e^+ \gamma) < 1.2 \times 10^{-11}$ (MEGA experiment, 1999)
- New experiment with a sensitivity of $\text{BR} \sim 10^{-14}$ planned at PSI



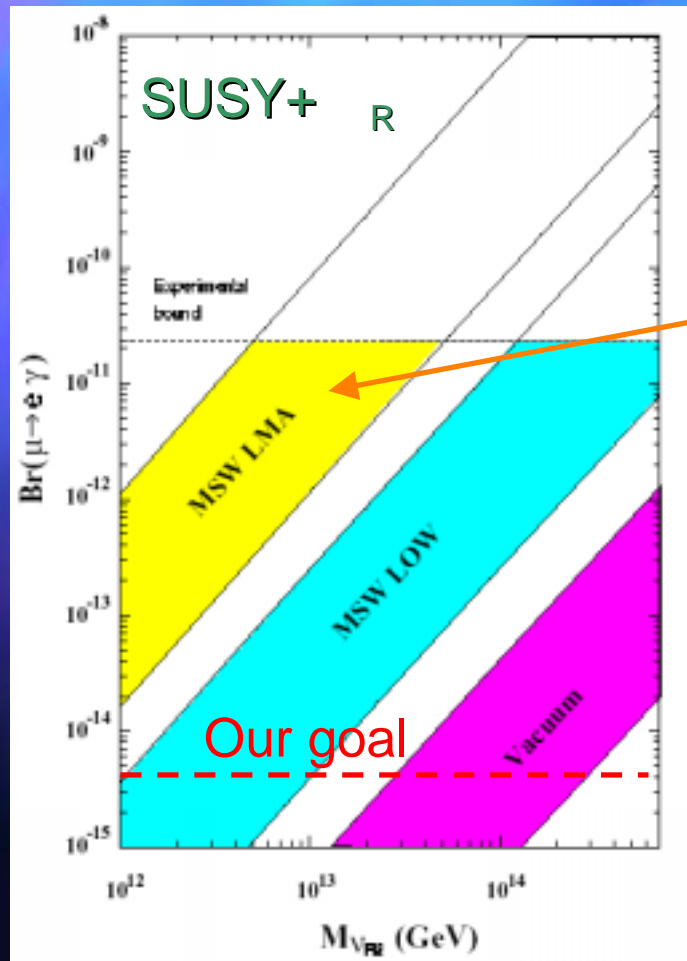
Physics Motivations



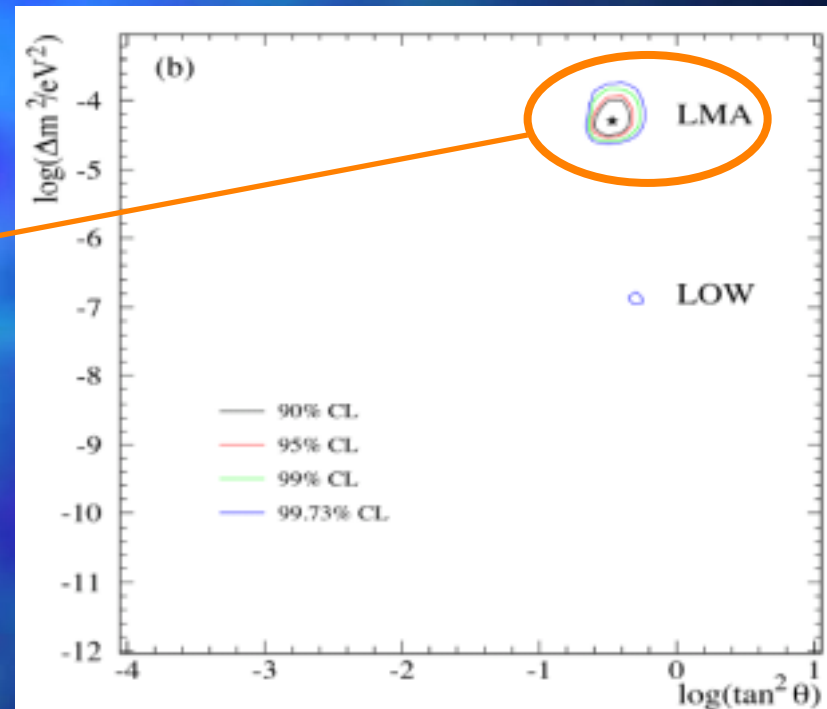
J. Hisano et al.,
Phys. Lett. B391 (1997) 341

- SU(5) SUSY-GUT predicts $\text{BR}(\mu \rightarrow e\gamma) = 10^{-15} - 10^{-13}$
(SO(10) SUSY-GUT: even larger value $10^{-13} - 10^{-11}$)
- Small $\tan\beta$ excluded by LEP SUSY search

Physics Motivations, cont'd



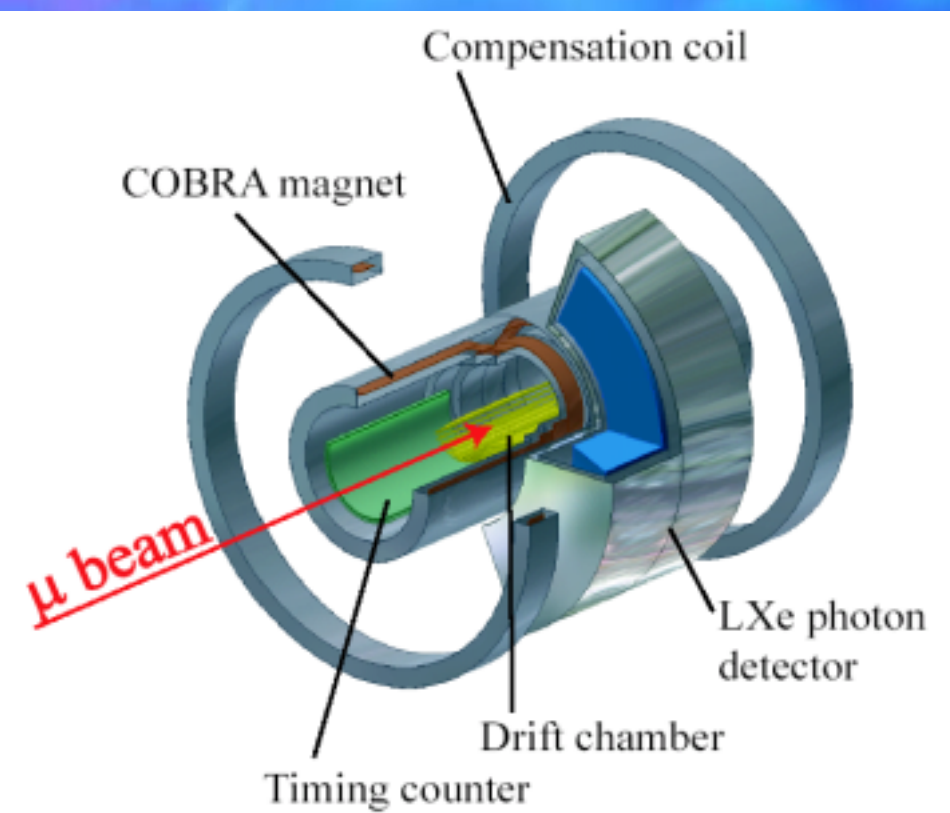
After the recent SNO measurements...



SNO collaboration, Q.R.Ahmad *et al.*, PRL89(2002)010302

- Solar ν meas. strongly favor the LMA.
- Large $\tan\beta \rightarrow$ large $\mu e \gamma$ rate

MEG Detector



- Liquid xenon photon detector
- Positron spectrometer with gradient magnetic field (COBRA spectrometer)
- World's most intense DC muon beam at PSI
- Sensitivity down to $BR \sim 10^{-14}$
- Engineering/physics run will start in 2004

Sensitivity and Background

- Single event sensitivity

$$N_{\mu} = 1 \times 10^8 / \text{sec}, T = 2.2 \times 10^7 \text{sec}, \Omega / 4\pi = 0.09, \varepsilon_{\gamma} = 0.7, \varepsilon_e = 0.95$$

$$\rightarrow \text{BR}(\mu^+ \rightarrow e^+ \gamma) \sim 0.94 \times 10^{-14}$$

- Major backgrounds

- Accidental Coincidence

Michel decay ($\mu^+ \rightarrow e^+ \bar{\nu}_e \nu_{\mu}$)
+ random

$$B_{\text{accidental}} \sim 5 \times 10^{-15}$$

- Radiative muon decays

$$\mu^+ \rightarrow e^+ \nu_e \bar{\nu}_{\mu}$$

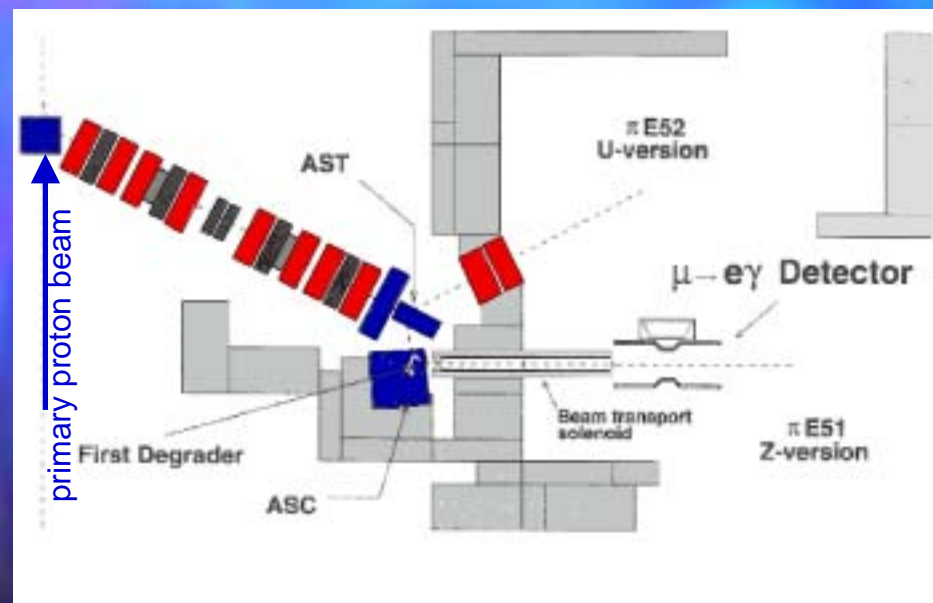
$$B_{\text{prompt}} \sim 10^{-17}$$

Proposed detector performance

E_e	0.7% (FWHM)
E	1.4 – 2.0 % (FWHM)
θ_e	12 – 14 mrad(FWHM)
t_e	0.15 nsec (FWHM)

These values could be changed according to the actually achieved performance of the detector.

Beam Line



- DC muon beam rate above $10^8 \mu/s$ at π E5 beam line
- Two beam branches (“U” and “Z”)
- Comparative study of the branches is in progress.
- Positron contamination can be reduced by:
 - (1) Combination of an energy degrader and a magnetic selection
 - (2) Wien filter

Condition	“Z”-branch	“U”-branch
No degrader, transmitted to zone	$3.6 \times 10^8 \mu^+/s$ $6.0 \times 10^8 e^+/s$	$3.5 \times 10^8 \mu^+/s$ $1.6 \times 10^9 e^+/s$
Degradator at final focus	$2.0 \times 10^8 \mu^+/s$	$3.2 \times 10^7 \mu^+/s$
m/e ratio at Muon Peak	9	16.5

Decision on the choice of the beam branch will be made after the beam tests

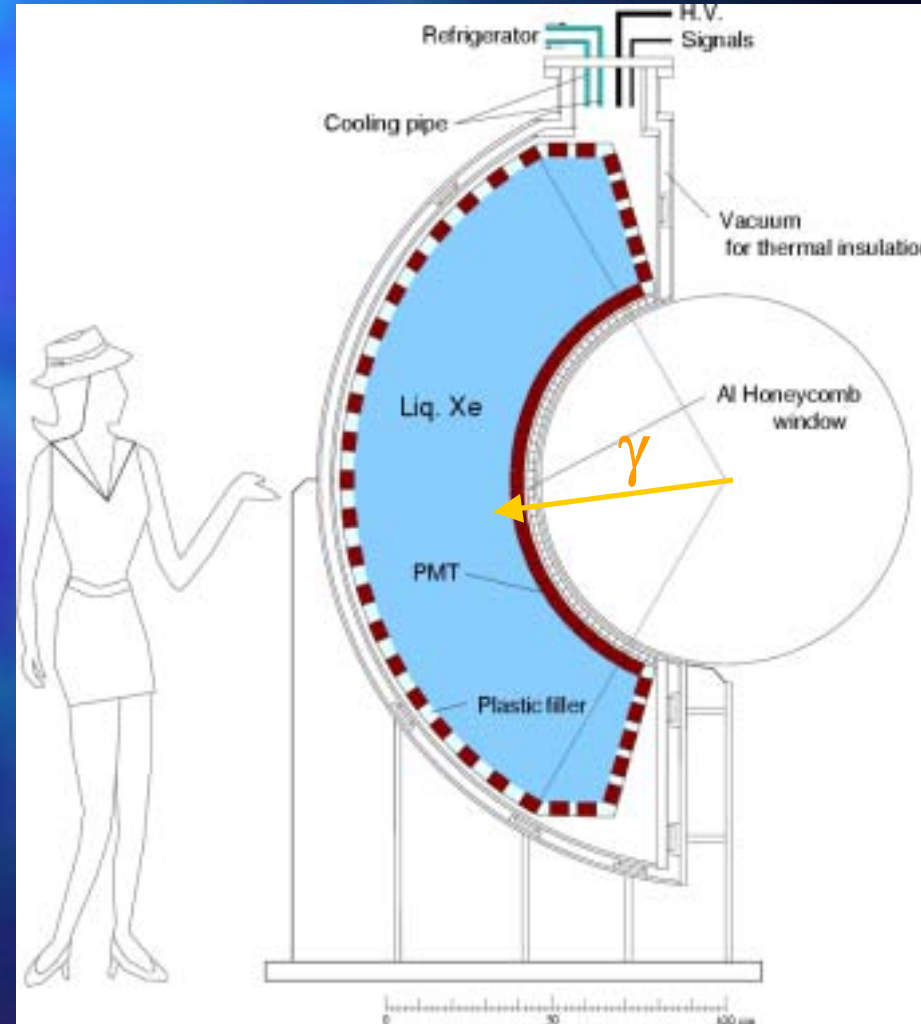
with “U”-branch in Aug.2002 and with “Z”-branch in Nov.2002

Liquid Xenon Photon Detector

- High light yield (75% of NaI(Tl))
- Fast signals
 - avoid accidental pileups
- Spatially uniform response
 - no need for segmentation

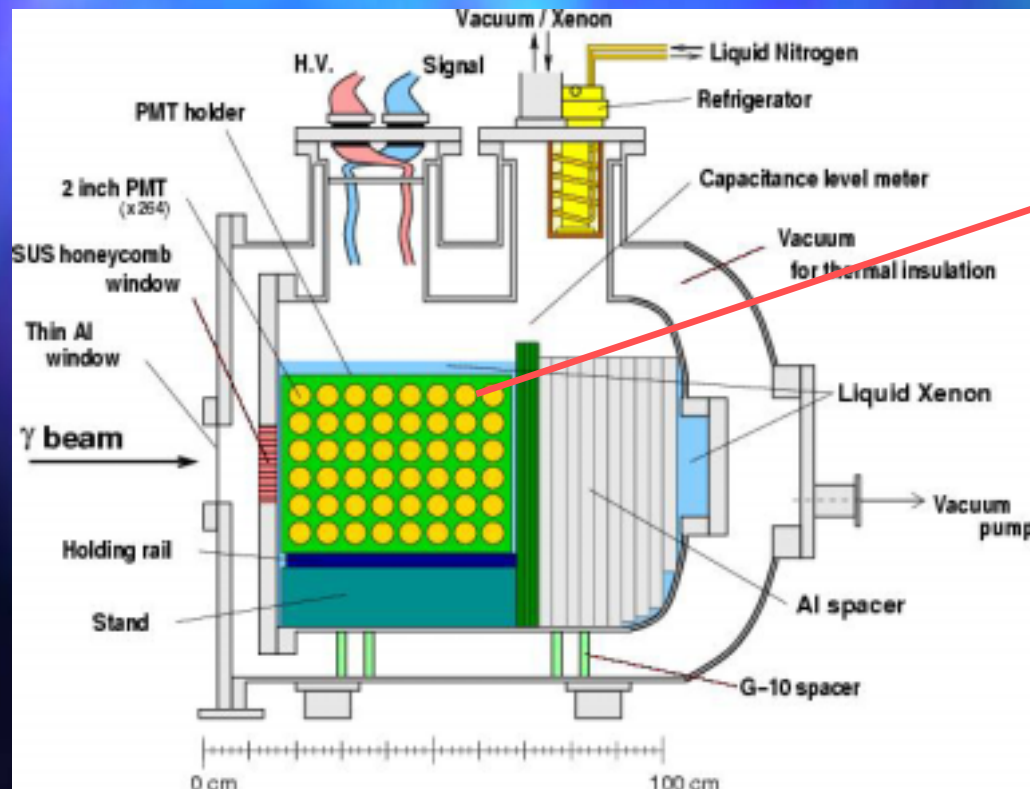
Current design

- Active volume of LXe: ~800 liter
- Scintillation light is collected by ~800 PMTs immersed in LXe
- Compact PMT with metal channel dynode structure and quartz window
(Hamamatsu R6041Q)



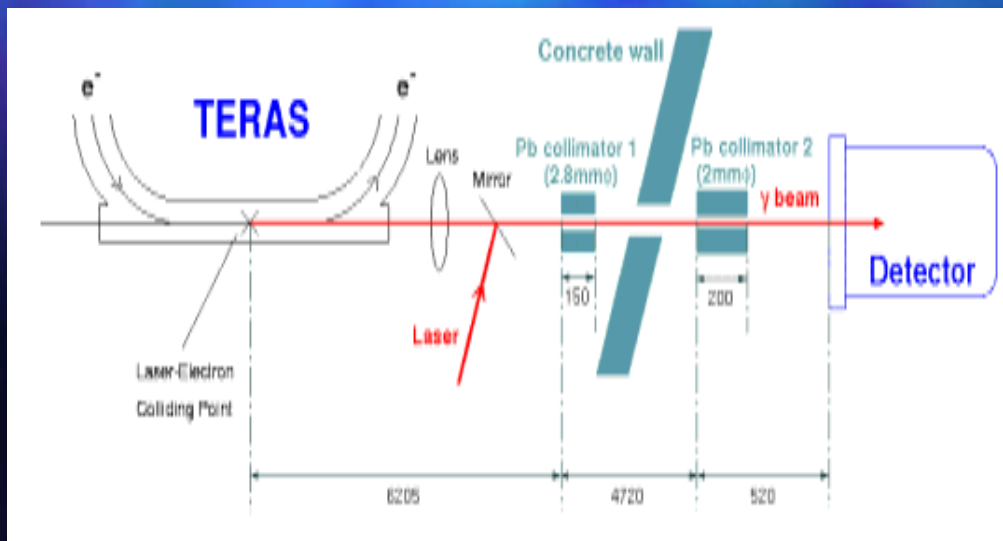
Photon Detector Prototype

- A total of 120 liter liquid xenon (**active volume of 69 liter**)
- Viewed by **240 PMTs**
- Large enough to test with $\sim 50\text{MeV } \gamma$
- LEDs and α sources (^{241}Am) implemented for calibration

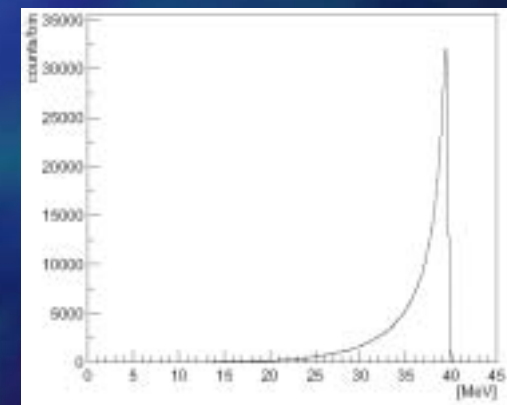


Gamma Beam Tests

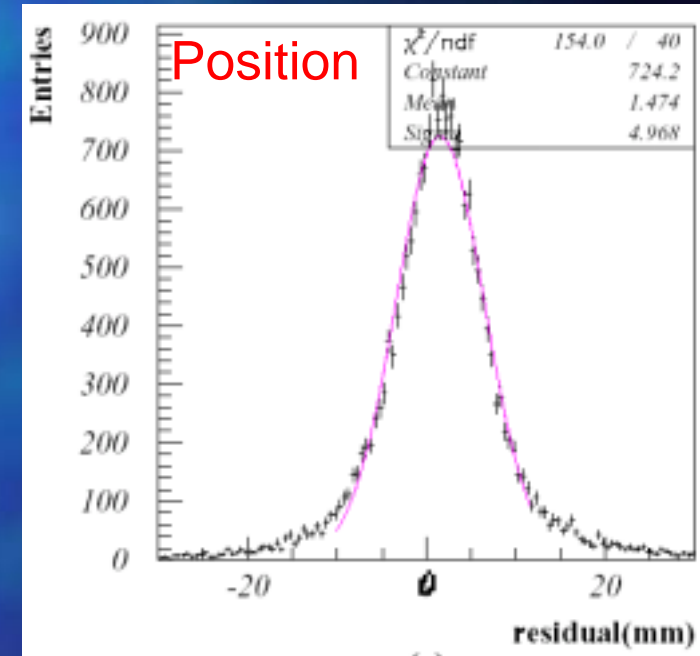
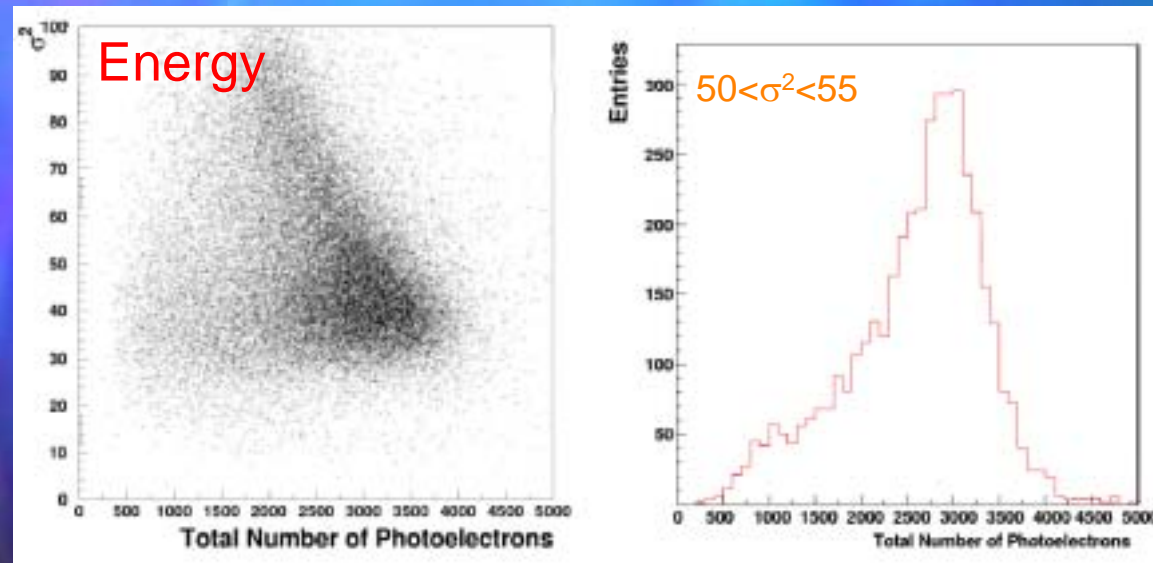
- Performance test of large prototype using high-energy gamma rays
- Laser Compton backscattering facility at TERAS electron storage ring of AIST, Tsukuba, Japan
- Gamma-ray beam with energy up to 40MeV
- Energy resolution evaluated by spread of Compton edge
- Position reconstructed by PMT output distribution with proper collimator
- Timing reconstructed by averaging arrival time
- **Beam test in Feb. 2002**



Energy spectrum of gamma beam with 1mm ϕ collimator (simulation)



Beam test in Feb. 2002



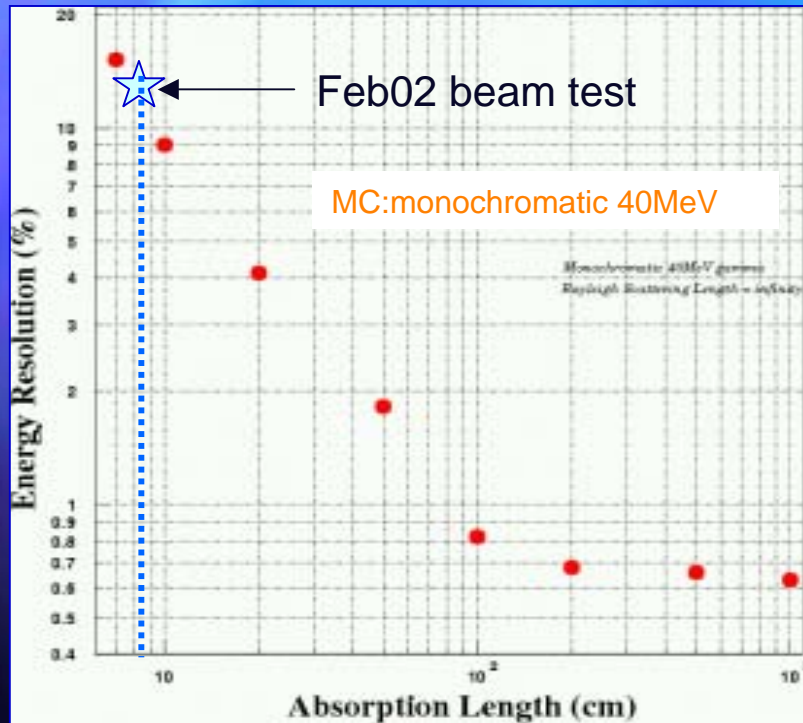
σ^2 : conversion depth parameter

- Observed amount of light from 40MeV γ is smaller than expected. (~10%)
- Strong correlation between the conversion depth and N_{pe}
- Worse position resolution than expected

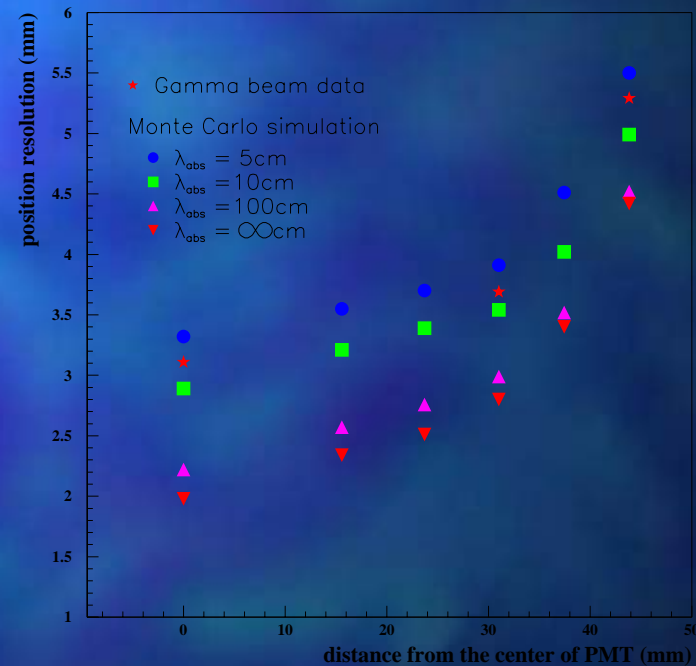
→ can be explained by strong light absorption in LXe

MC Predictions with Absorption

Energy resolution



Position resolution

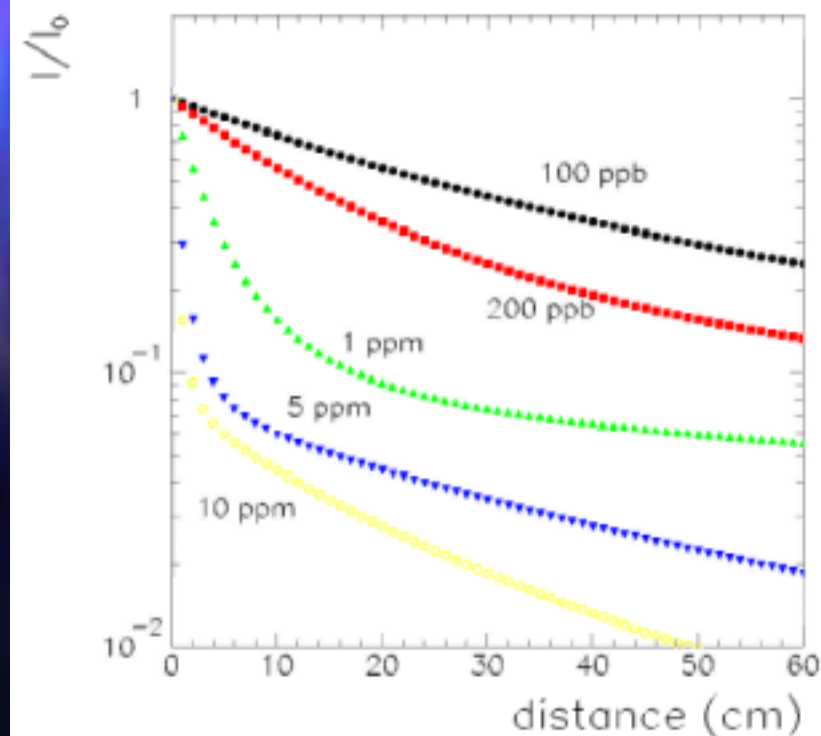


- MC predictions indicate $\lambda_{\text{abs}} < 10\text{cm}$ in gamma beam test in Feb. 2002
- We need $\lambda_{\text{abs}} > 100\text{cm}$ at least for an energy resolution of a few % order

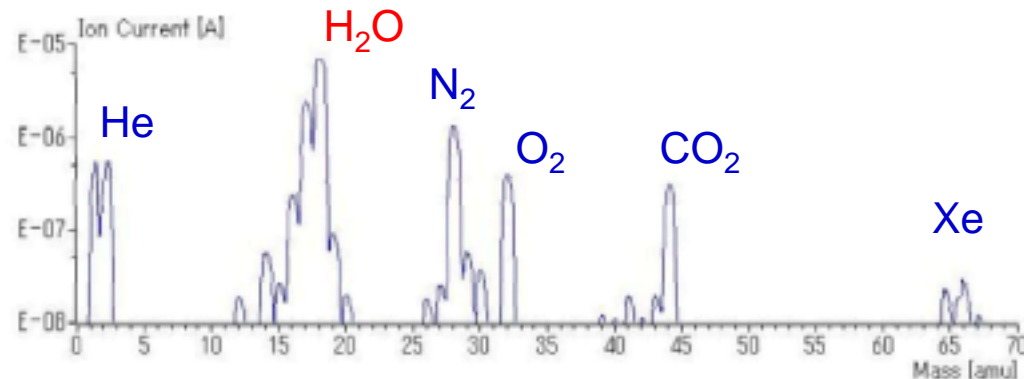
Light Absorption in LXe

H_2O , C_2H_4 , NH_3 , O_2 can strongly absorb 175nm scintillation light from LXe → Contaminations in LXe?

Absorption in H_2O -doped Xe

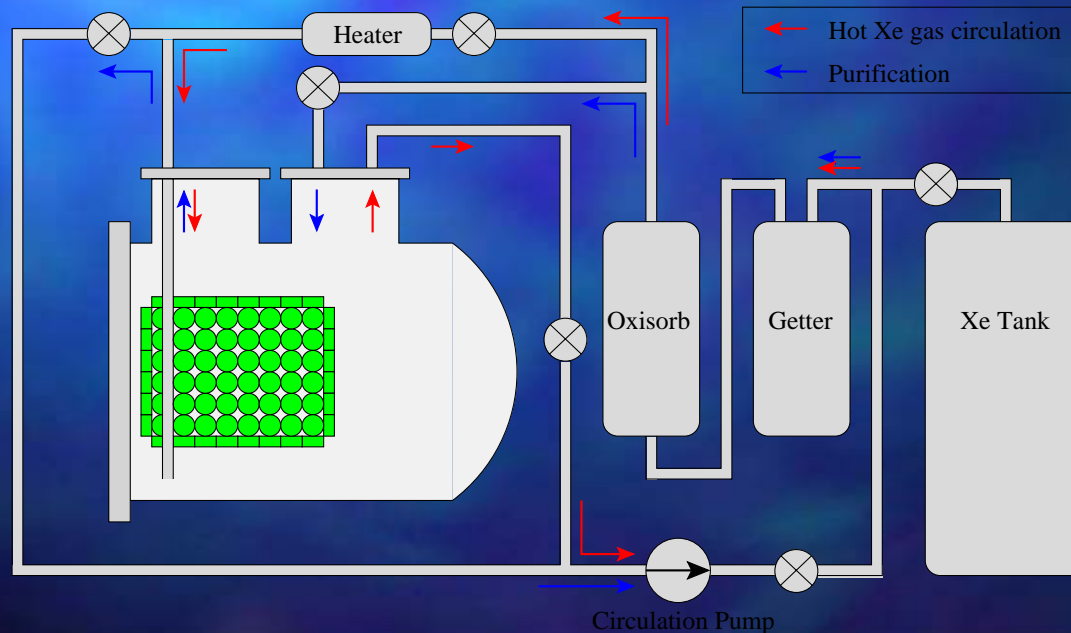


Mass spectrum for the remaining gas in the detector vessel

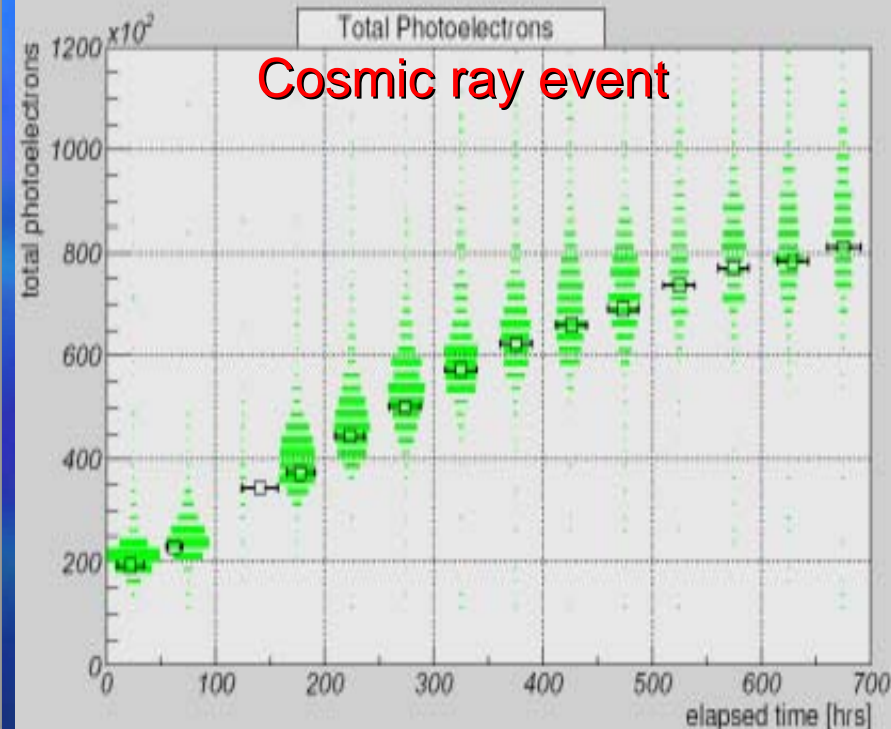
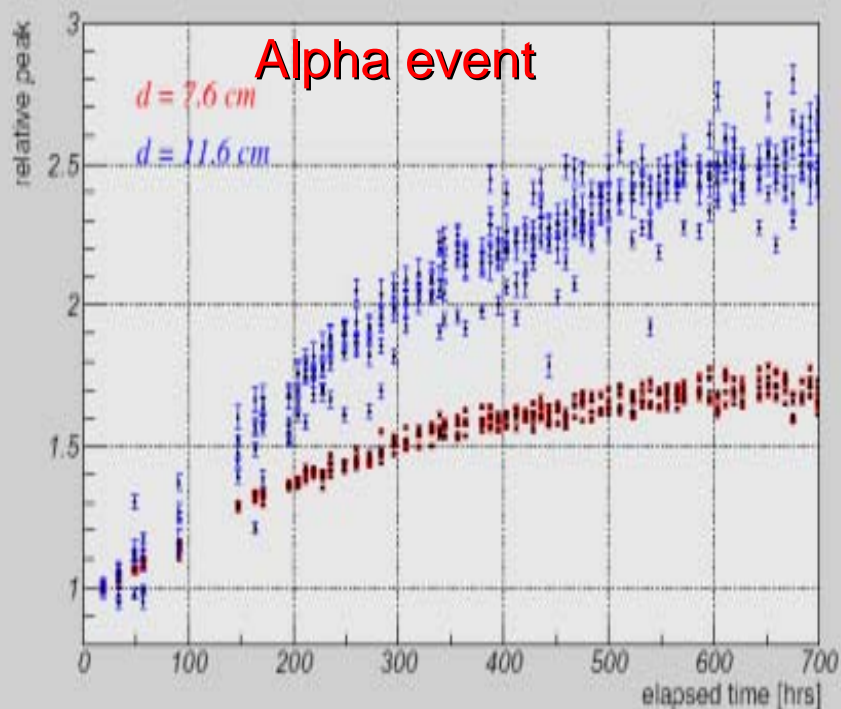


Purification

- New circulatory purification system is installed after the beam test in Feb.2002.
- Xenon vapor is purified in Zr-V-Fe getter and Oxisorb filter and recondensed by the refrigerator and LN₂ during the operation of the detector
- Circulation speed 10-12cc liq./minute



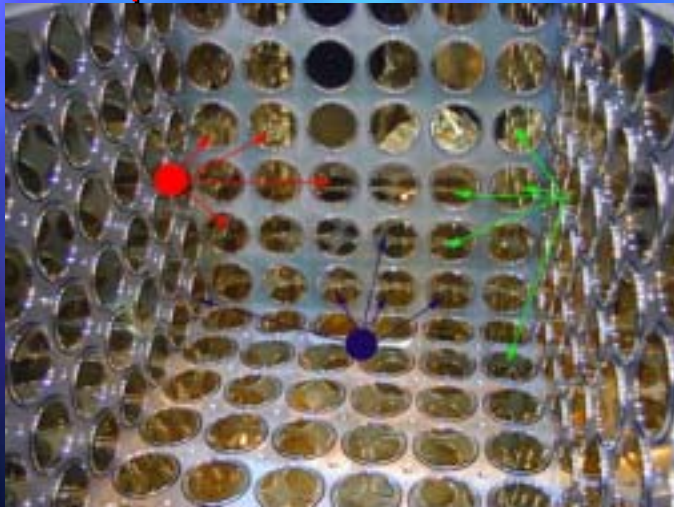
Improvement of Light Yield



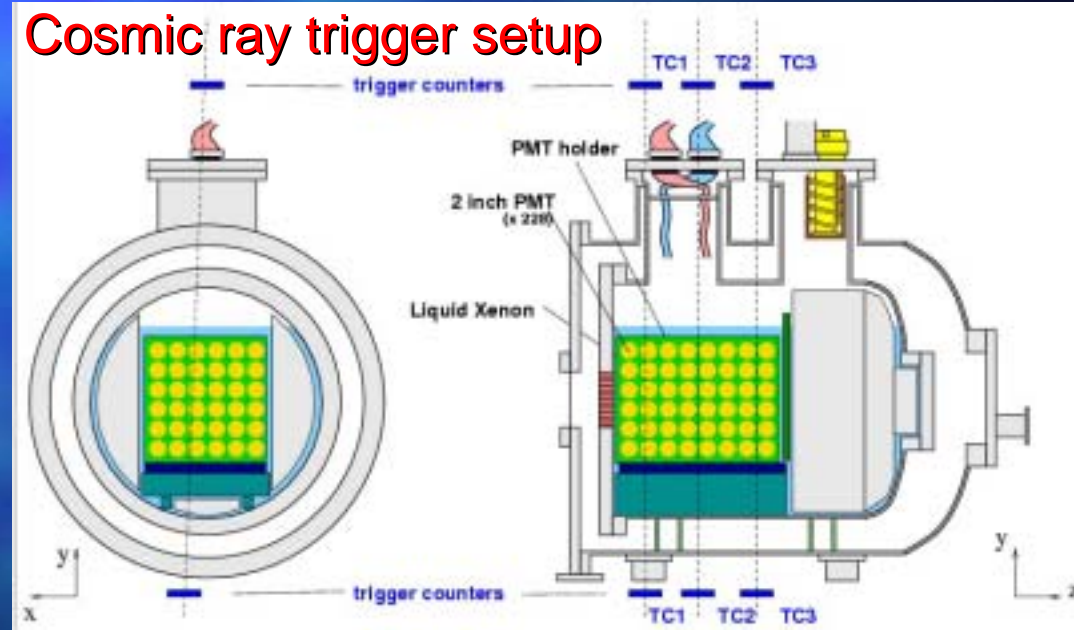
Absorption Length Estimation

Absorption length is estimated by seeing the absorption of the light from the alpha source event and cosmic ray event.

4 x alpha source inside

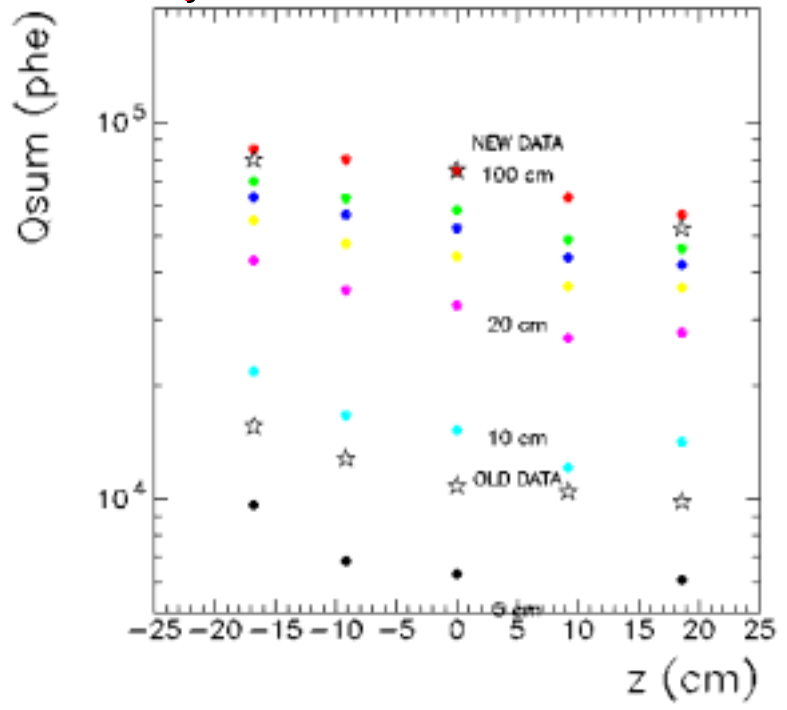


Cosmic ray trigger setup

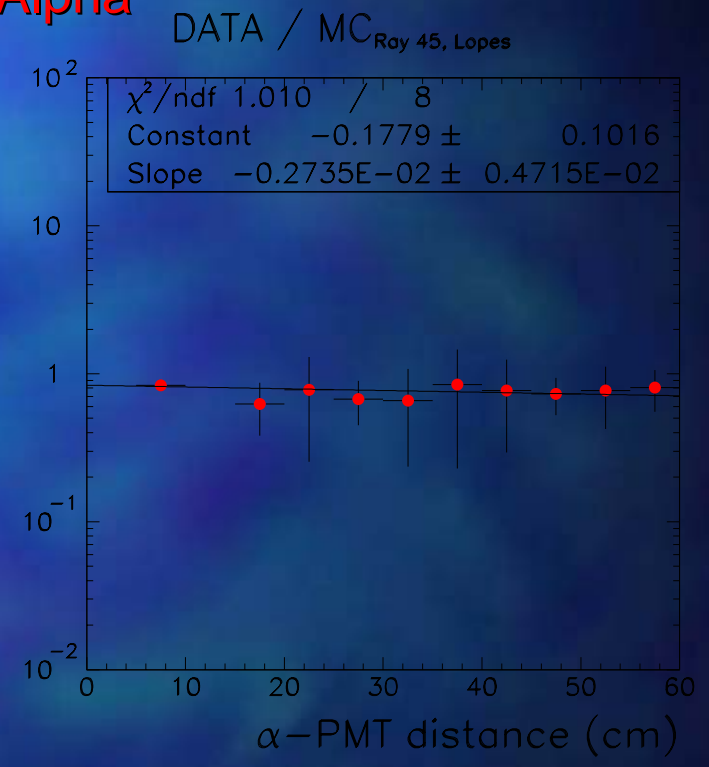


Absorption Length Estimation, cont'd

Cosmic ray



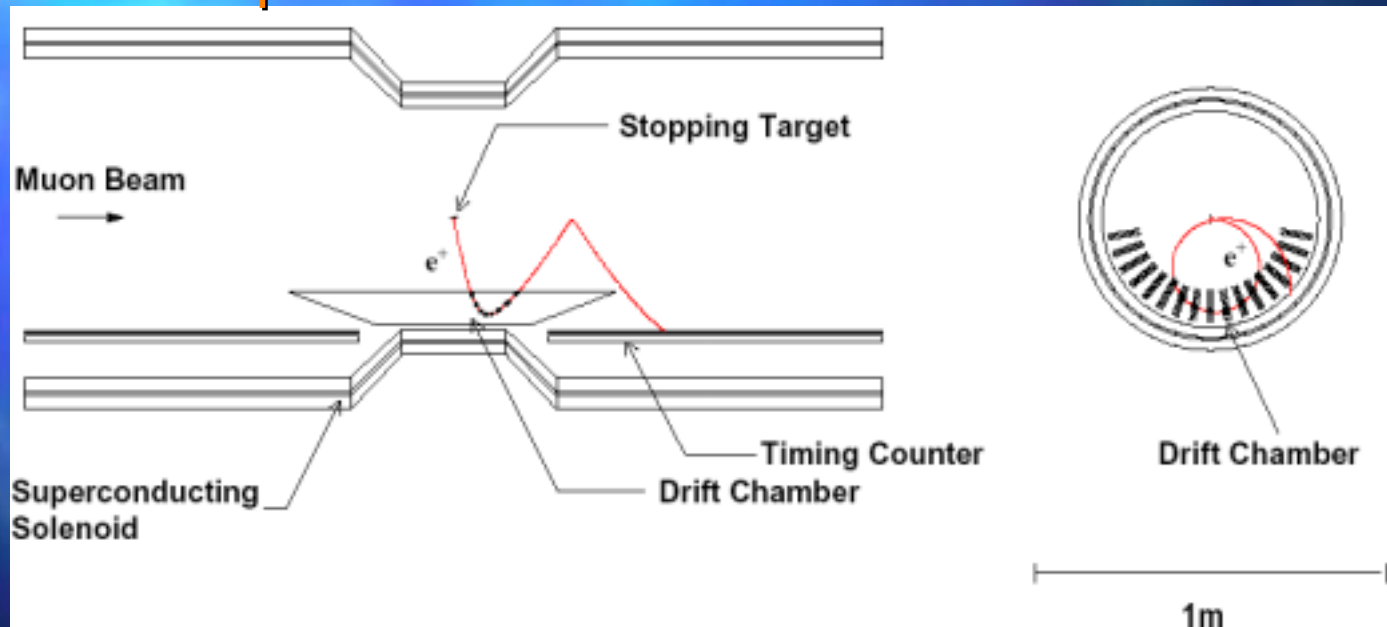
Alpha



Both measurements (CR and α) indicate $\lambda_{abs} \sim 100\text{cm}$ after the purification

Positron Spectrometer

COBRA spectrometer

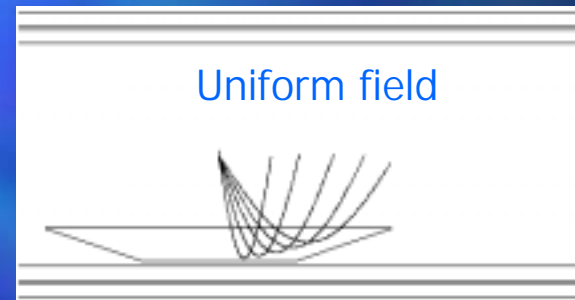
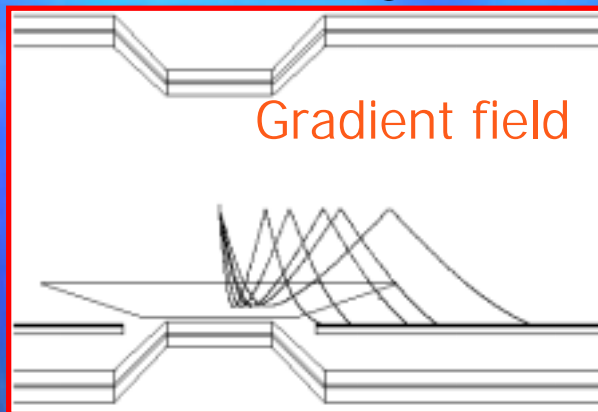


- Thin superconducting magnet designed to form gradient magnetic field
- Drift chamber for positron tracking
- Scintillation counters for timing measurement

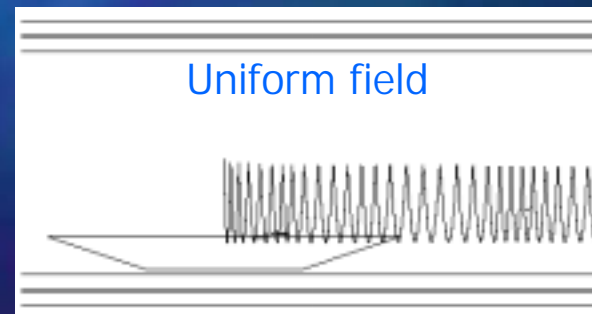
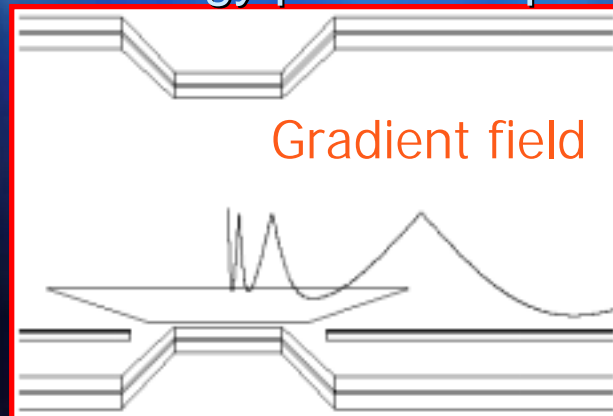
Concept of COBRA Spectrometer

COBRA : COnstant Bending RAdius

- Constant bending radius independent of emission angles

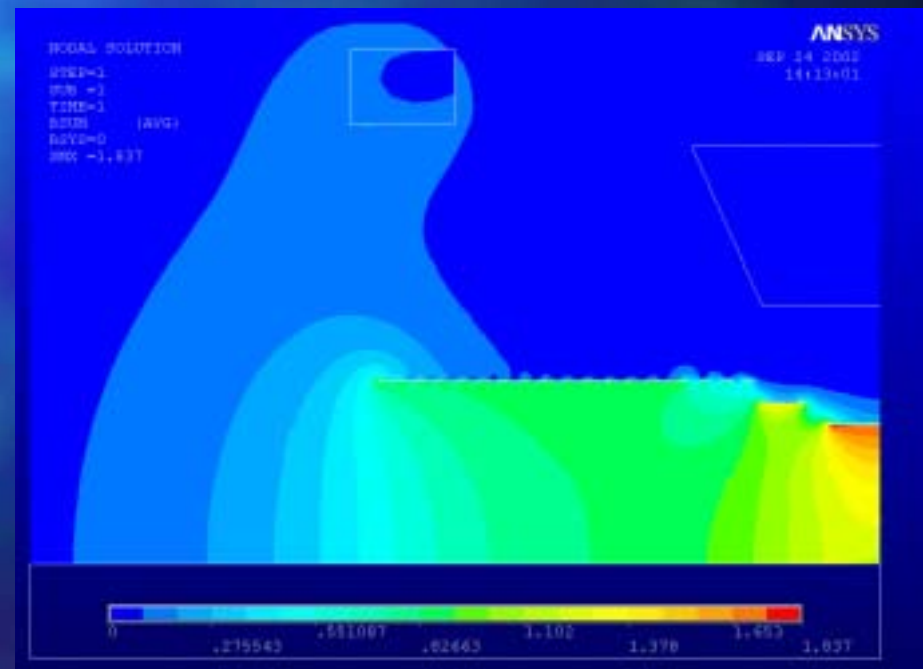
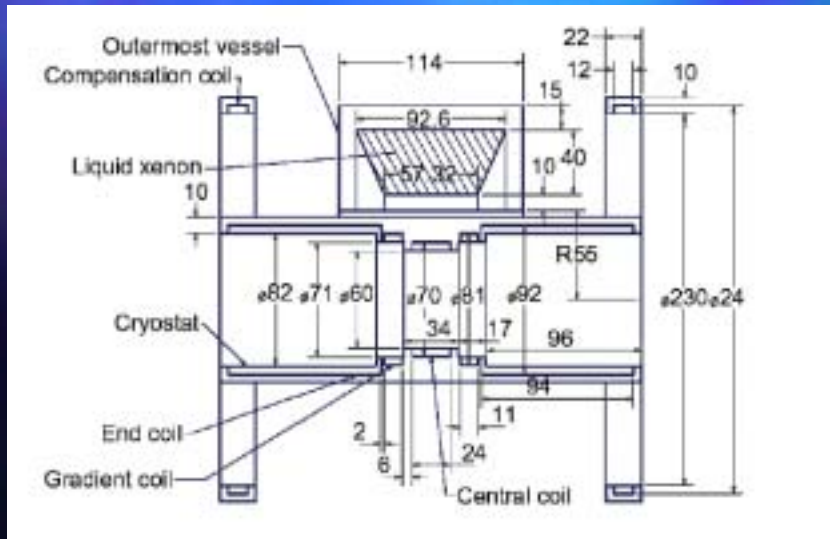


- Low energy positrons quickly swept out



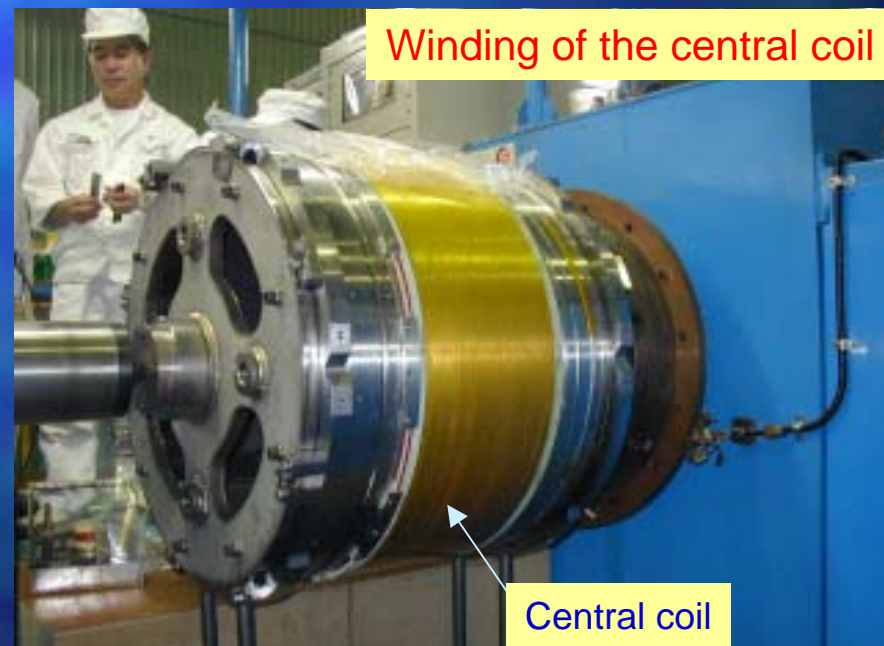
Magnet

- Five coils with three different diameter to form **gradient field**
- $B_c = 1.26\text{T}$, $B_{z=1.25\text{m}} = 0.49\text{T}$ @ operating current = 359A
- **Compensation coils to suppress the residual field around the LXe detector down to ~50Gauss**
- High-strength aluminum stabilized superconductor
→ thin superconducting coil: $0.2X_0$

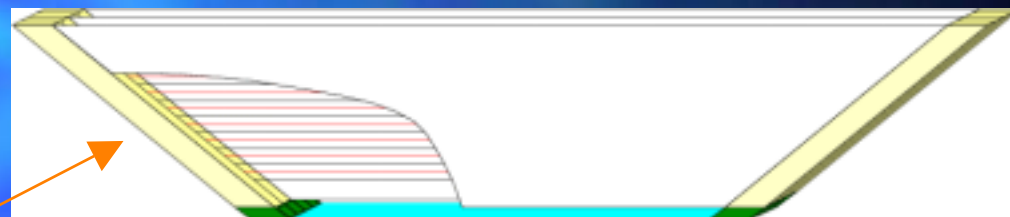
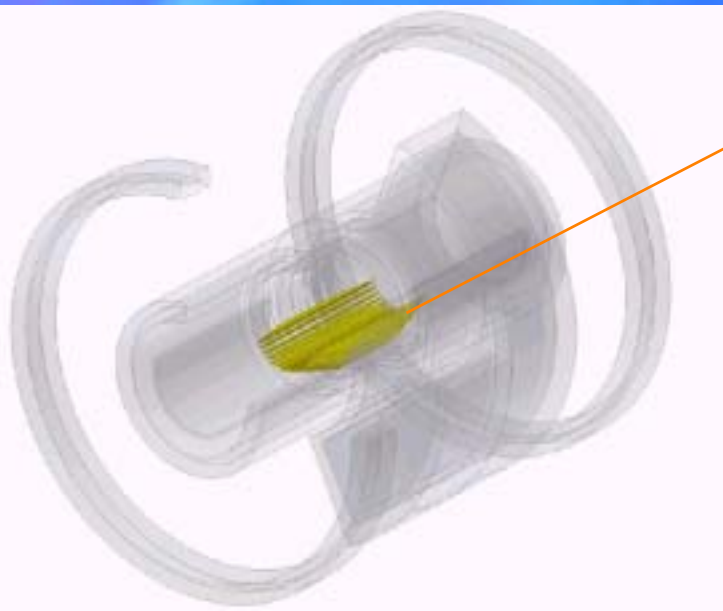


Construction of the Magnet

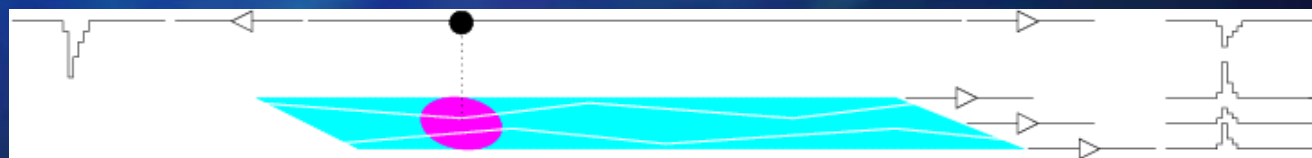
- Magnet design was finalized after detailed mechanical calculations and related experimental tests.
- Winding of the cable is in progress @ Toshiba.
- Excitation test for the central part of the magnet will be performed in October 2002.



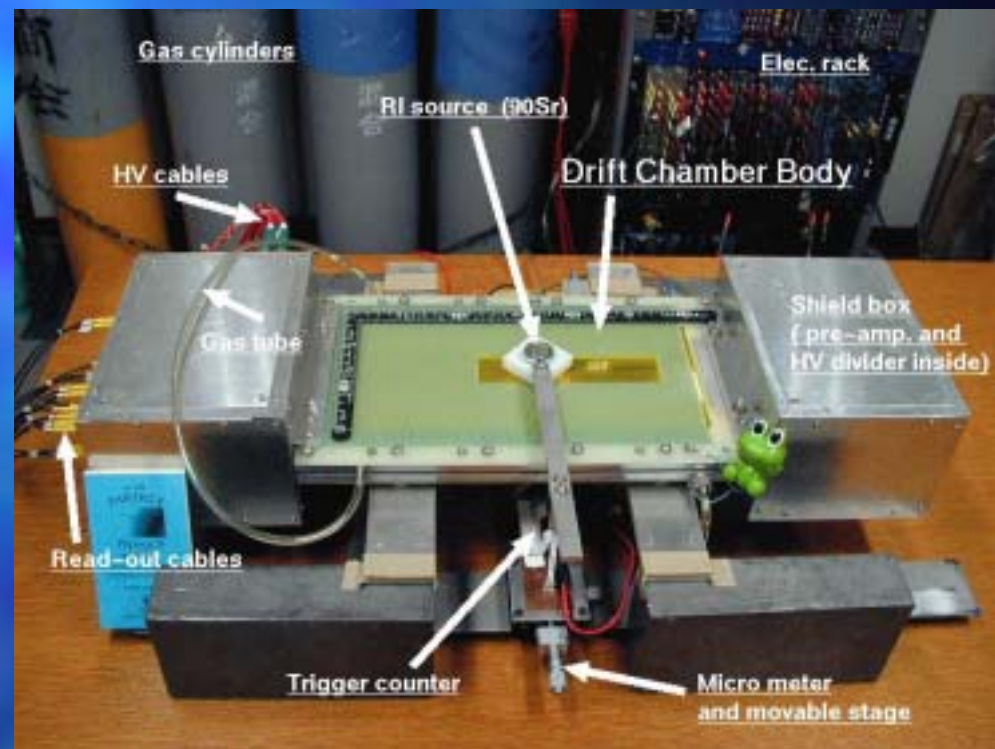
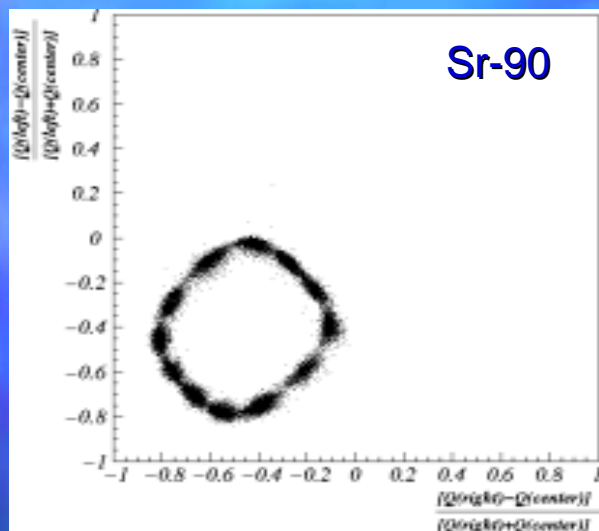
Positron Tracker



- 17 chamber sectors aligned radially with 10° intervals
- Two staggered arrays of drift cells
- Chamber gas: He-C₂H₆ mixture
- Vernier pattern on the cathode foil to determine z-position



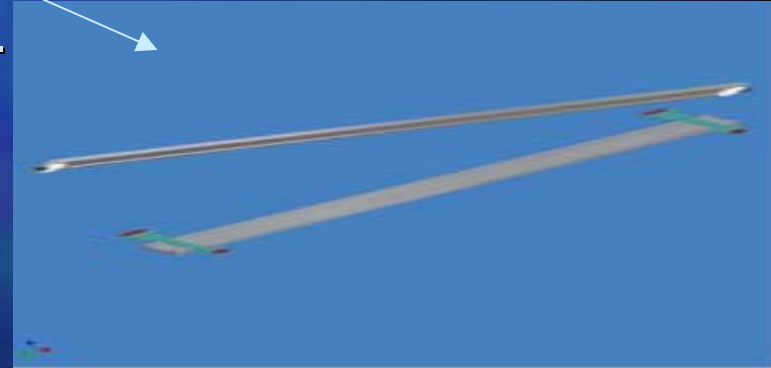
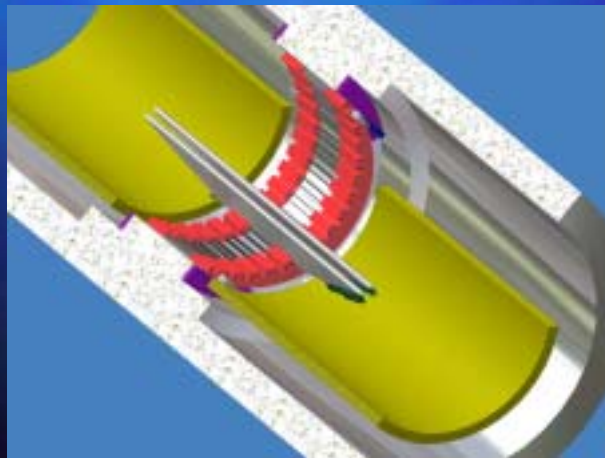
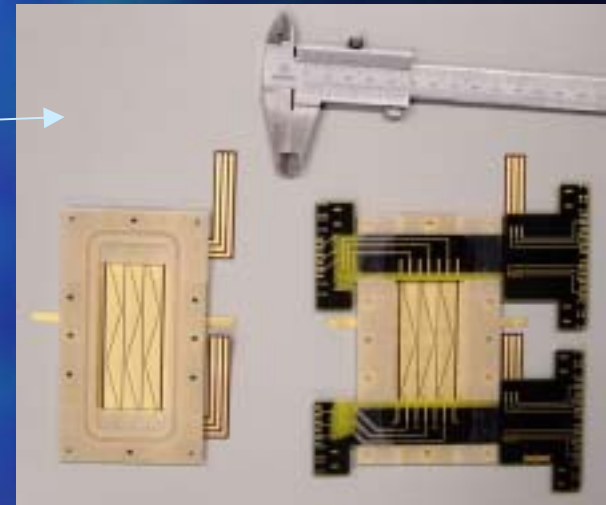
First Prototype of the Chamber



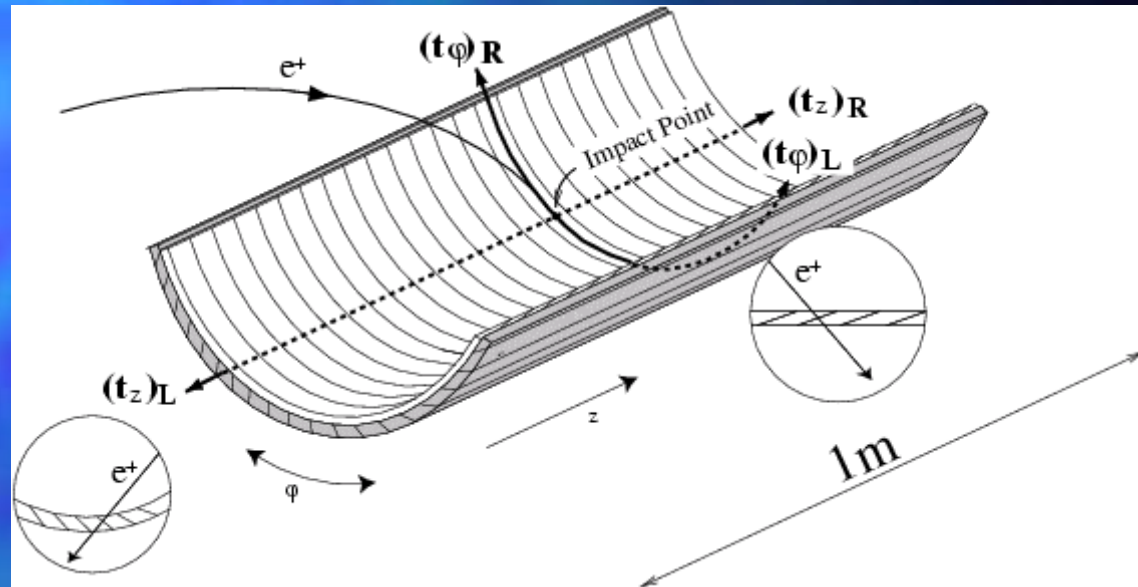
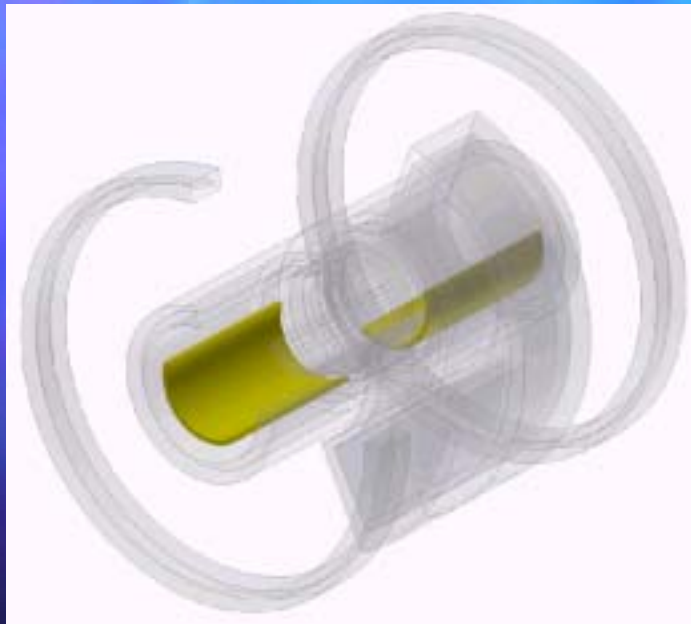
Resolution(σ)	
Drift time measurement	100-150 μm
Vernier cathod measurement	425 μm
Charge division measurement	2cm
Drift velocity and drift time	4-12ns

Chambers System R&D in PSI

- Two prototypes are under construction at PSI.
 - **“Double cathode” test chamber**
 - Two separated double-strip cathodes for each chamber layer
 - homogeneous position sensitivity
 - Test in 1 Tesla magnetic field
 - **“Charge division” test chamber**
 - Charge division test
 - 1m-long W(330W/m) or Steel(1200W/m)
- Supporting system is also under development.



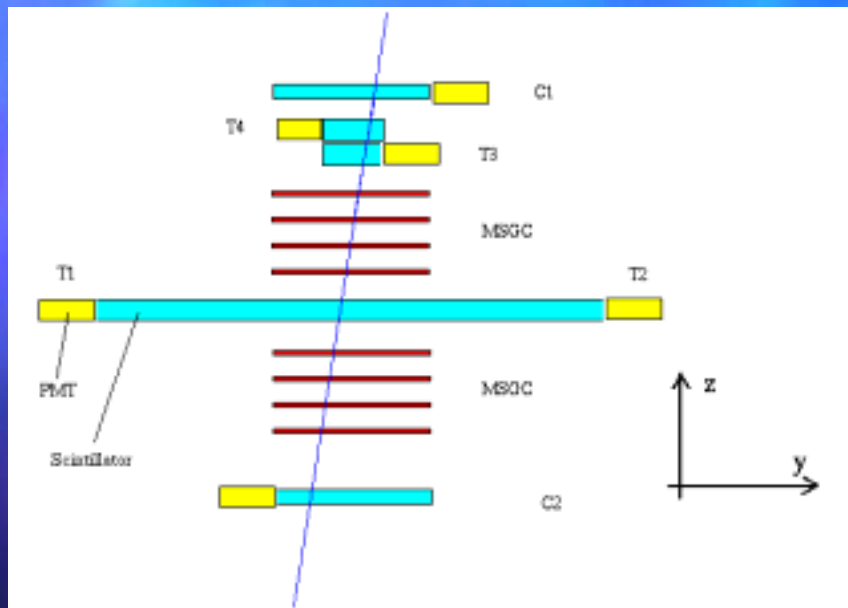
Timing Counter



- Two layers of scintillator hodoscopes placed at right angles with each other
 Outer: **timing measurement**
 Inner: **additional trigger information**
- Goal $\sigma_{\text{time}} \sim 50\text{psec}$

Timing Counter Prototype

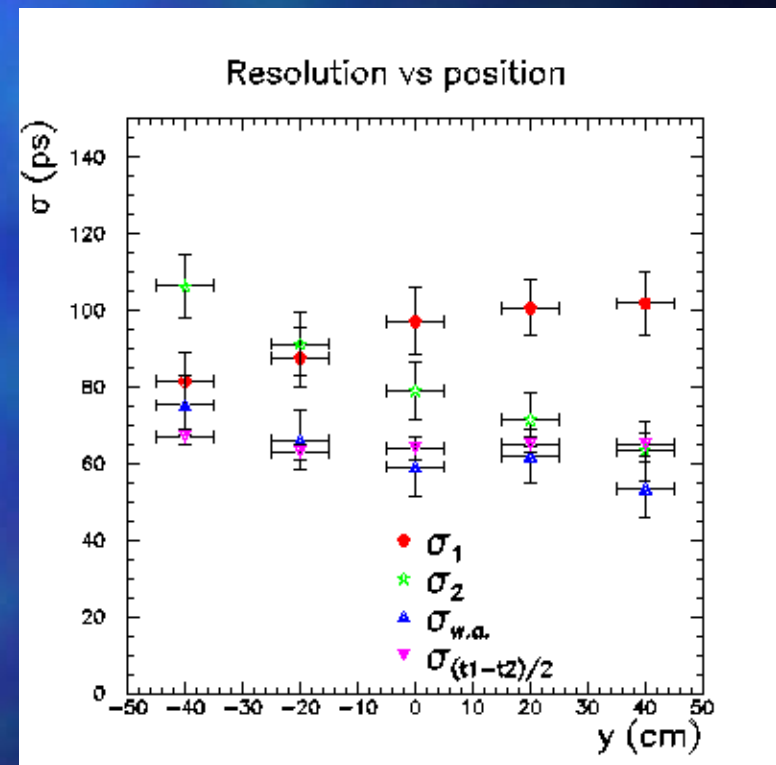
CORTES: Timing counter test facility with cosmic rays at INFN-Pisa



- Scintillator bar (5cm x 1cm x 100cm long)
- Telescope of 8 x MSGC
- Measured resolutions

$\sigma_{\text{time}} \sim 60 \text{ psec}$ independent of incident position

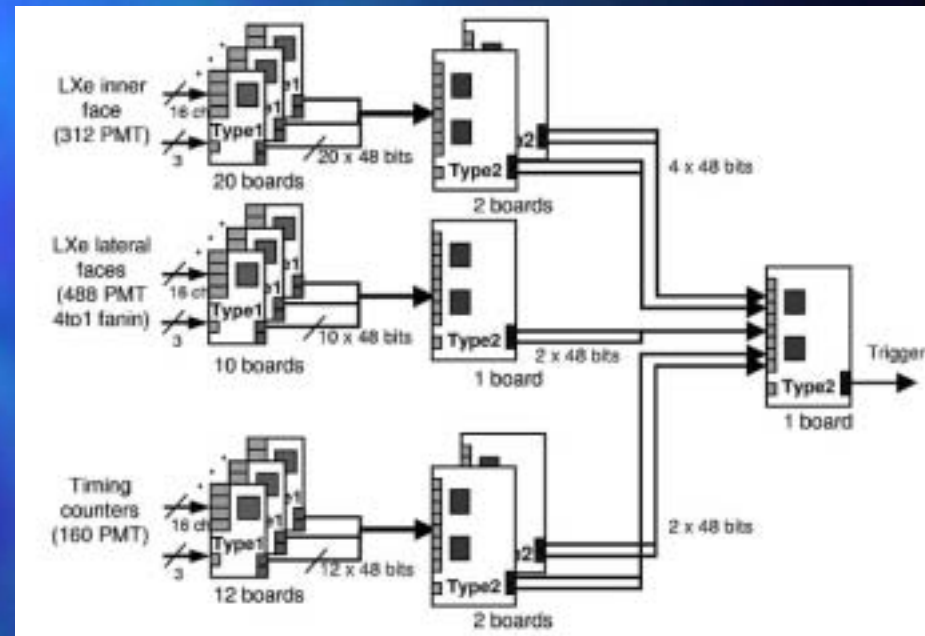
- σ_{time} improves as $\sim 1/\sqrt{N_{\text{pe}}}$ → use thicker counter $\sim 2 \text{ cm}$



Trigger Electronics

- Beam rate 10^8 s^{-1}
- Fast LXe energy sum $>45\text{MeV}$ $2 \times 10^3 \text{ s}^{-1}$
- γ interaction point
- e^+ hit point in timing counter
- Time correlation γ - e^+ 200 s^{-1}
- Angular correlation 20 s^{-1}

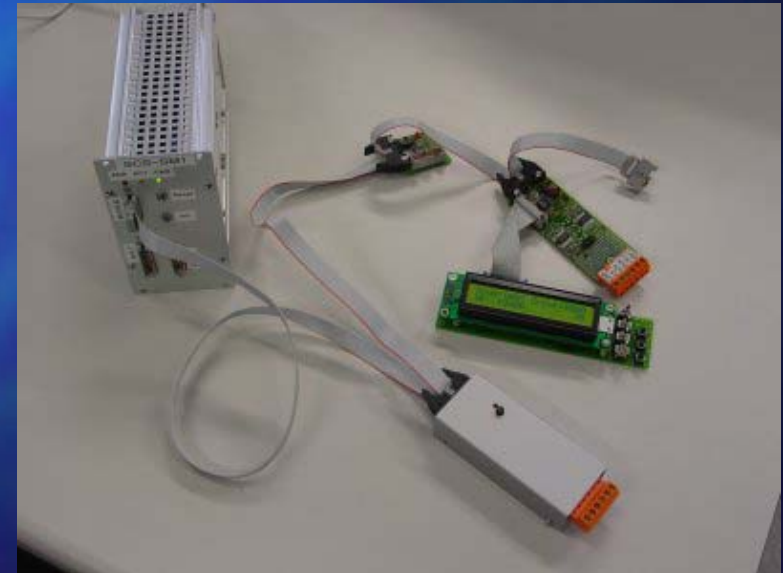
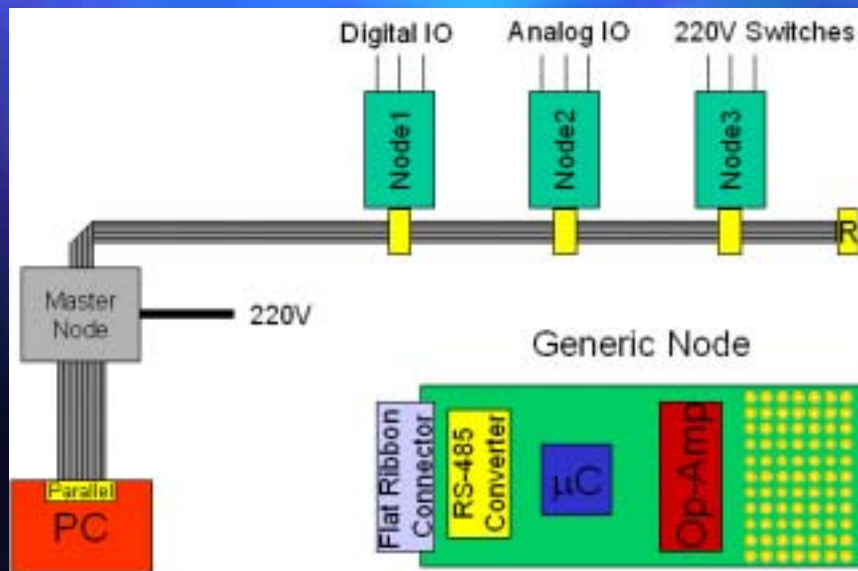
Trigger system structure



- Design and simulation of type1 board completed
- Prototype board delivered in Pisa by this fall

Slow Control

- New field bus system under development for a reliable control of cryogenics of LXe detector, superconducting magnet, high voltage supply
- Low cost (typ. 20 US\$ per node)
- Several prototypes have been built and tested at PSI
- See <http://midas.psi.ch/mscb>



Summary

- R&D work on the sub-detectors for the MEG experiment are going well.
- Performance of the LXe photon detector prototype is improving thanks to the improvement of the light yield.
- A beam test of the photon detector prototype with the purified xenon will be performed in Oct. 2002.
- Beam line tuning with the COBRA magnet and assembly of the sub-detectors will start in 2003.
- Engineering run will start in 2004.

Updated status can be seen at three mirrored sites:

<http://meg.icepp.s.u-tokyo.ac.jp/>

<http://meg.psi.ch/>

<http://meg.pi.infn.it/>